

# **The SXS Catalog of Simulations**

**Mike Boyle**

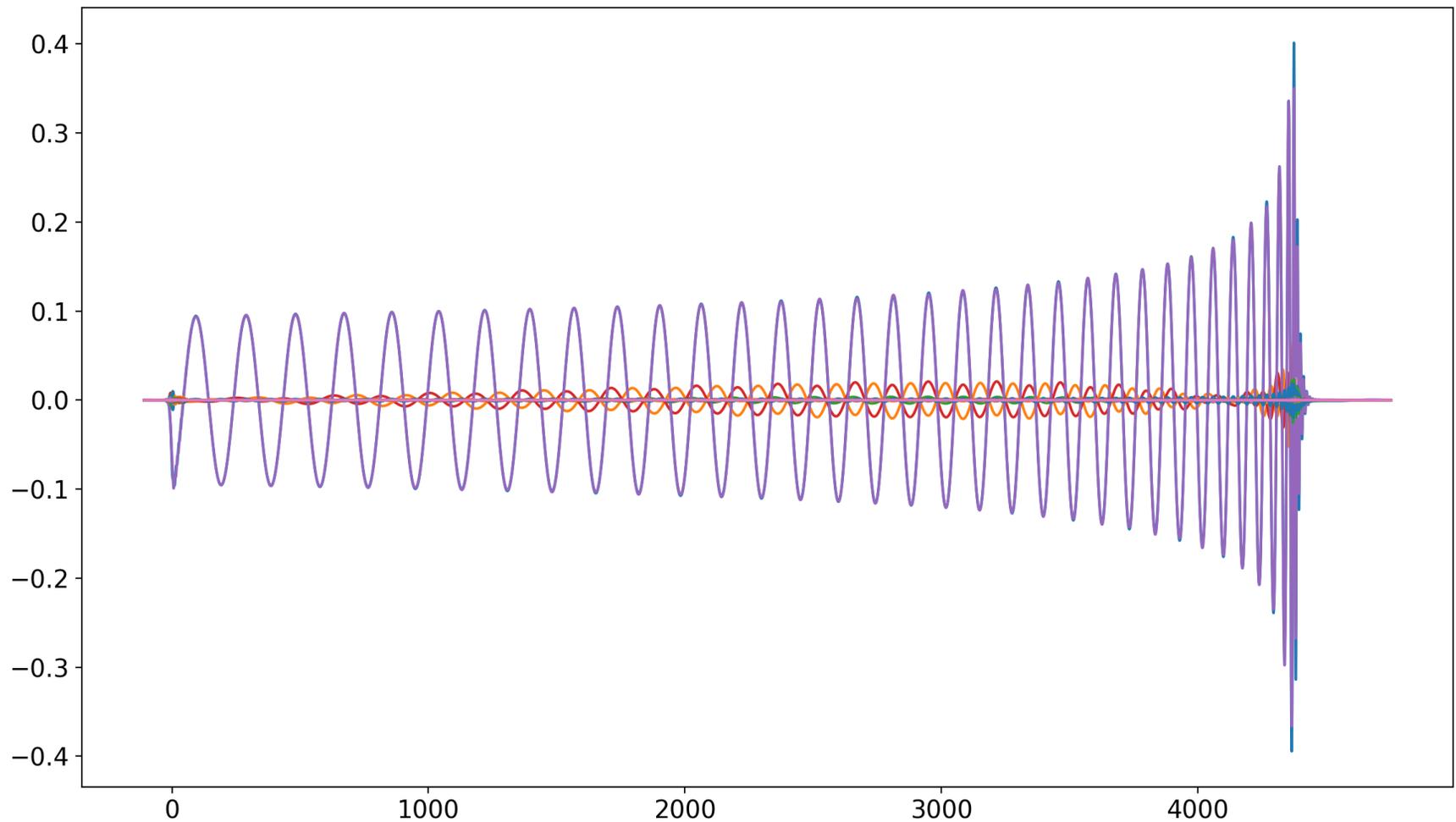
# Outline

- Simulations
  - SpEC
  - The numbers
  - Post-processing
- The data
  - What's included
  - How to get it
  - Formats
  - Problems
  - Solutions
- Software
  - Backwards compatible
  - Future proof

```
In [1]: import numpy as np
import matplotlib.pyplot as plt
import sxs

waveform = sxs.load("SXS:BBH:0123/Lev/rhOverM", extrapolation_order=2, download=True)
plt.plot(waveform.t, waveform.data.real);
```

Skipping download from 'https://data.black-holes.org/catalog.json' because local file is newer  
Found the following files to load from the SXS catalog:  
SXS:BBH:0123v4/Lev5/rhOverM\_Asymptotic\_GeometricUnits\_CoM.h5



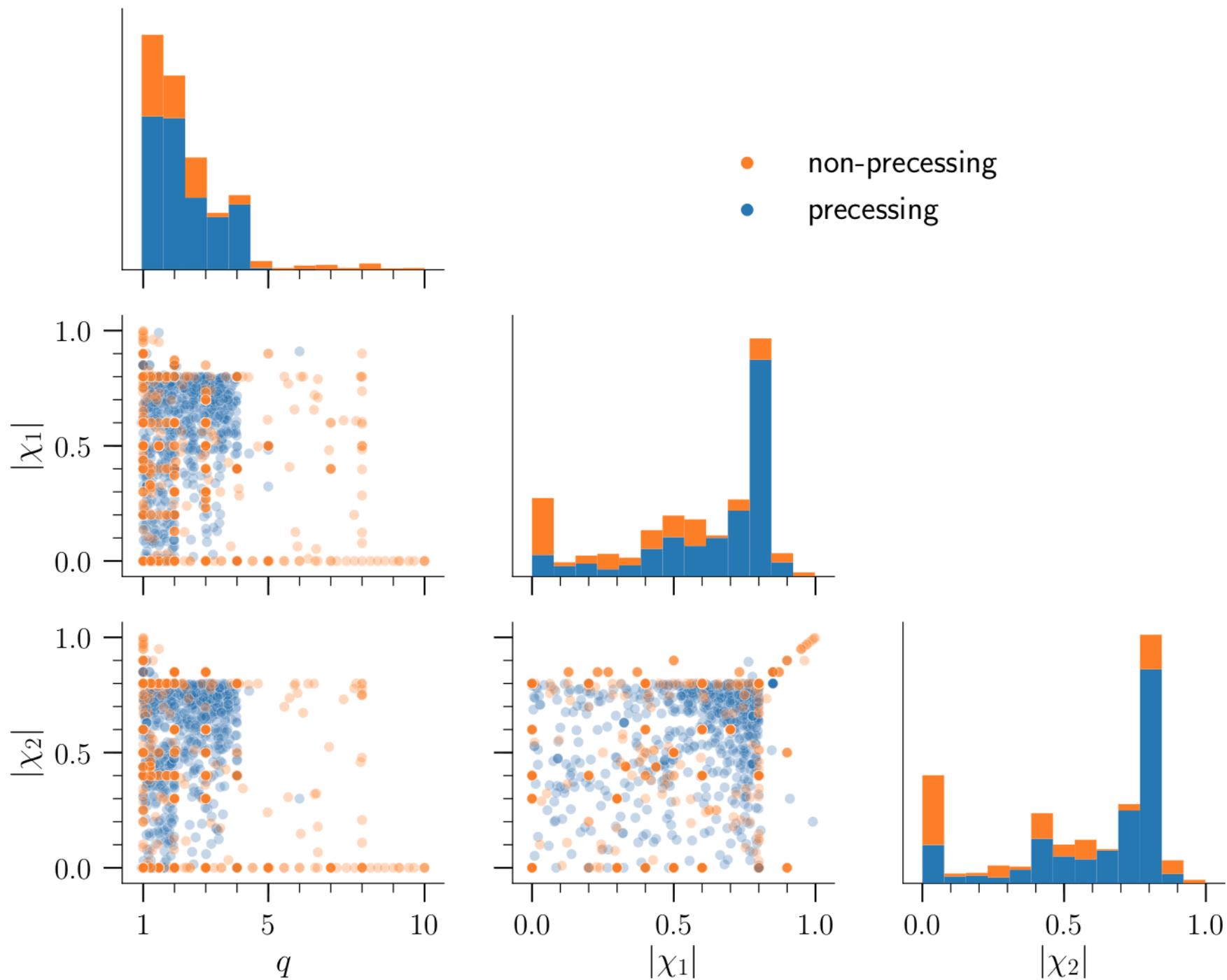
# SpEC

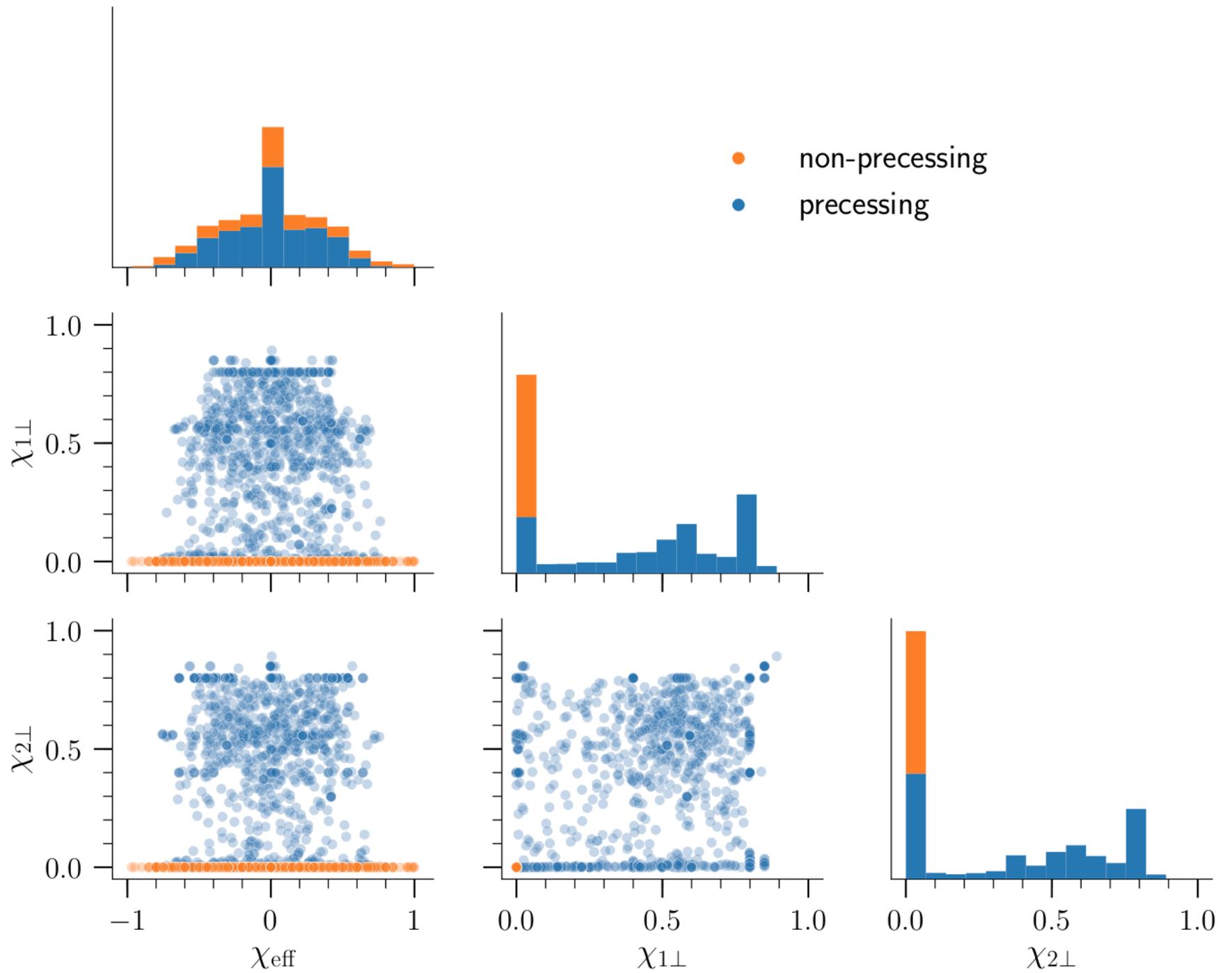
- Initial data solves extended conformal thin sandwich equations
- Iterate to reduce eccentricity if desired
- Multi-domain spectral methods
- First-order generalized harmonic with constraint damping
- Adaptive mesh refinement
- Constraint-preserving boundary conditions

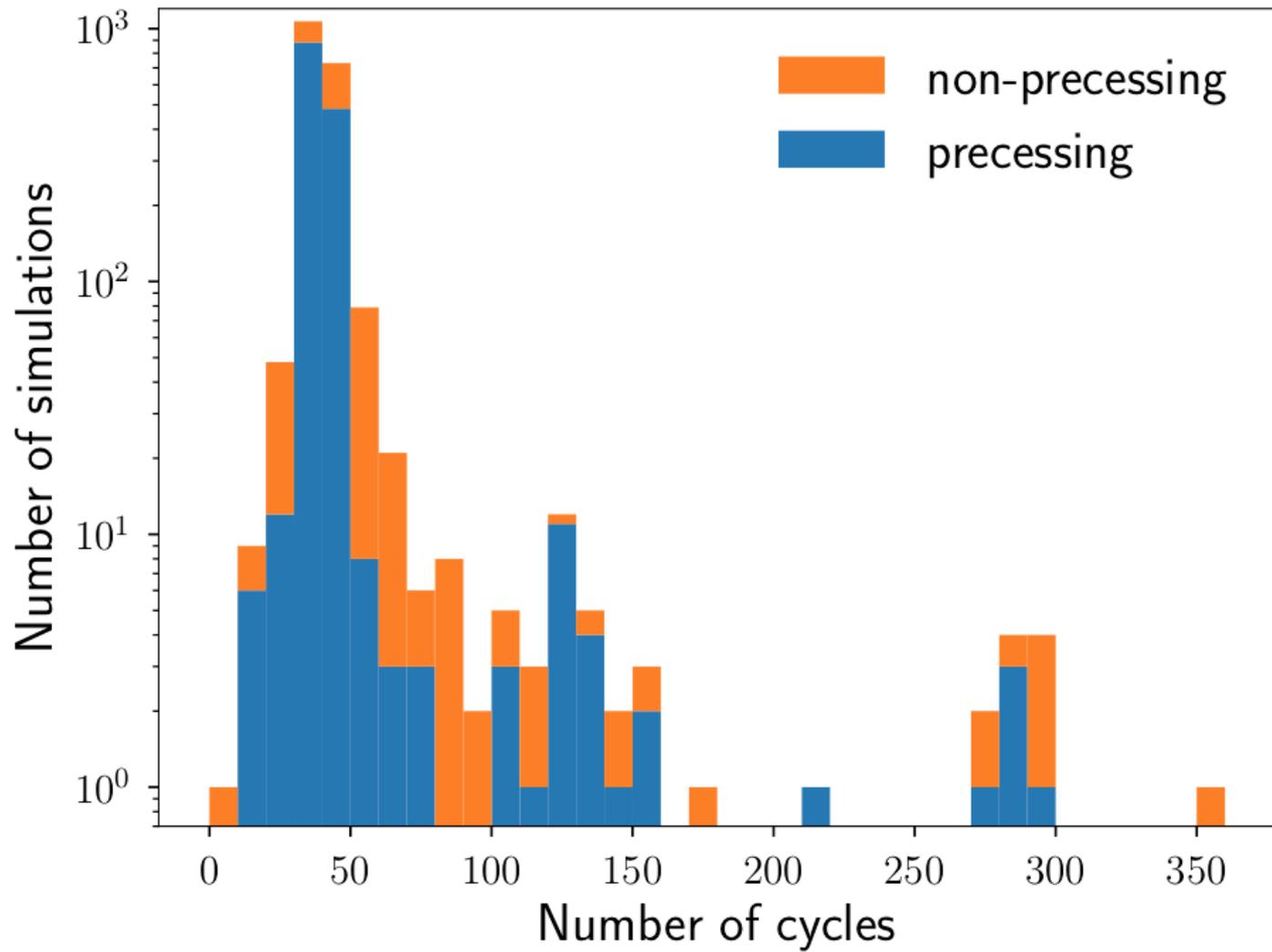
<https://arxiv.org/abs/1904.04831> (<https://arxiv.org/abs/1904.04831>)

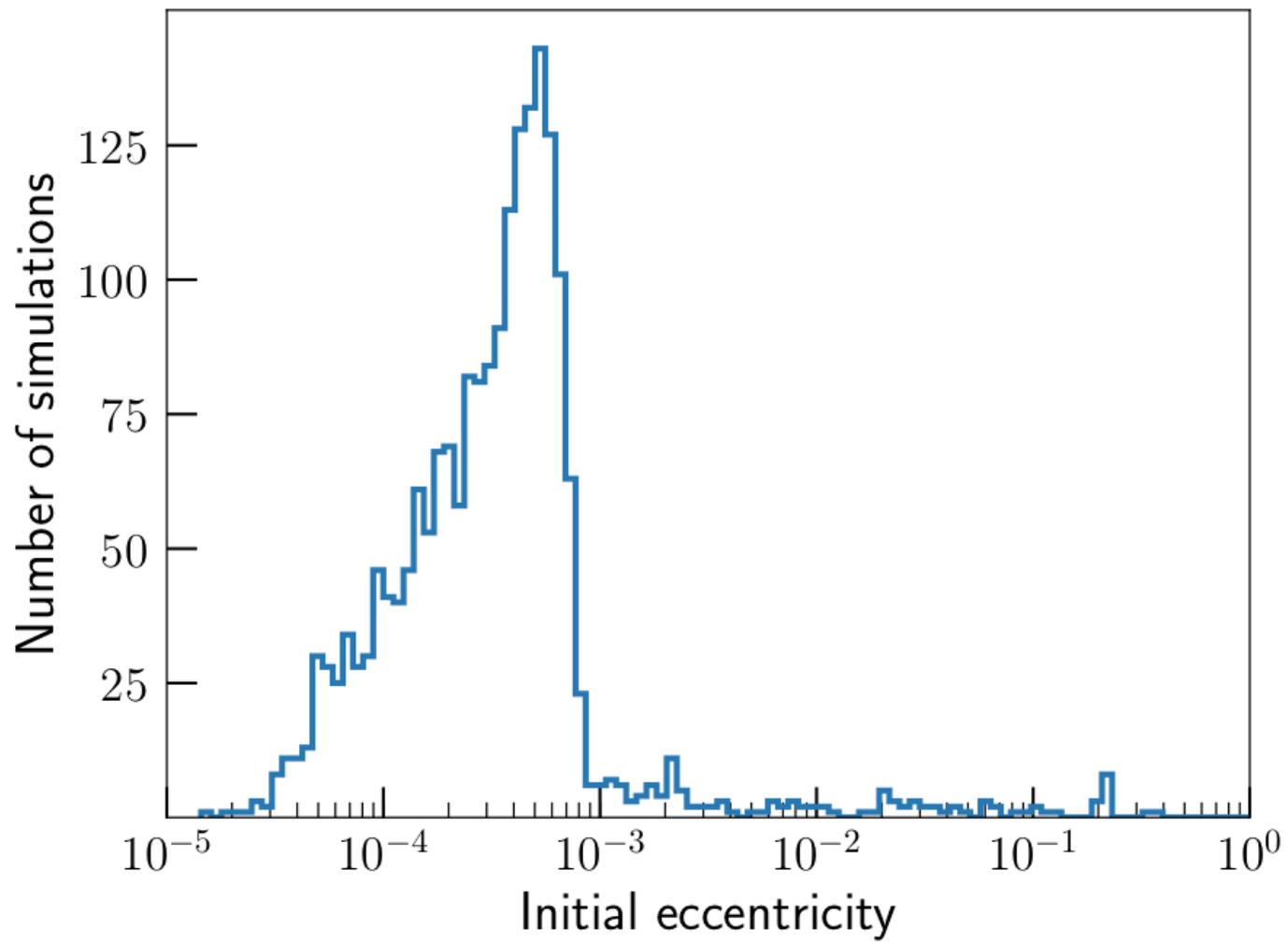
# Simulations

- 2 NSNS
- 7 BHNS
- 2019 BBH
  - mass ratios between 1 and 10
  - spin magnitudes up to 0.998
  - 1426 precessing simulations
  - typically ~19 orbits









# Data extracted during the runs

- Horizon information
  - Coordinate trajectories
  - Areal (irreducible) mass
  - Christodoulou mass
  - Spins
- On a series of  $\sim 24$  spheres between  $\sim 100M$  and outer boundary
  - RWZ  $h$  modes
  - $\Psi_4$  modes
  - Surface area
  - Average lapse
- (Data for CCE)

# Post-processing

- (CCE, soon)
- Extrapolation of  $\Psi_4$  and  $h$
- Center-of-mass correction

# Extrapolation

- Define corrected retarded time on each sphere

$$u_{i,j} = \int_0^{T_i} \sqrt{\frac{-1/g_j^{TT}(T)}{1 - 2M_{\text{ADM}}/R_j(T)}} dT - R_j(T_i) - 2M_{\text{ADM}} \ln \left[ \frac{R_j(T_i)}{2M_{\text{ADM}}} - 1 \right],$$

- Fit the waveforms with a simple

$$f(u, r) \approx \sum_{j=0}^N \frac{f_{(j)}(u)}{r^j}$$

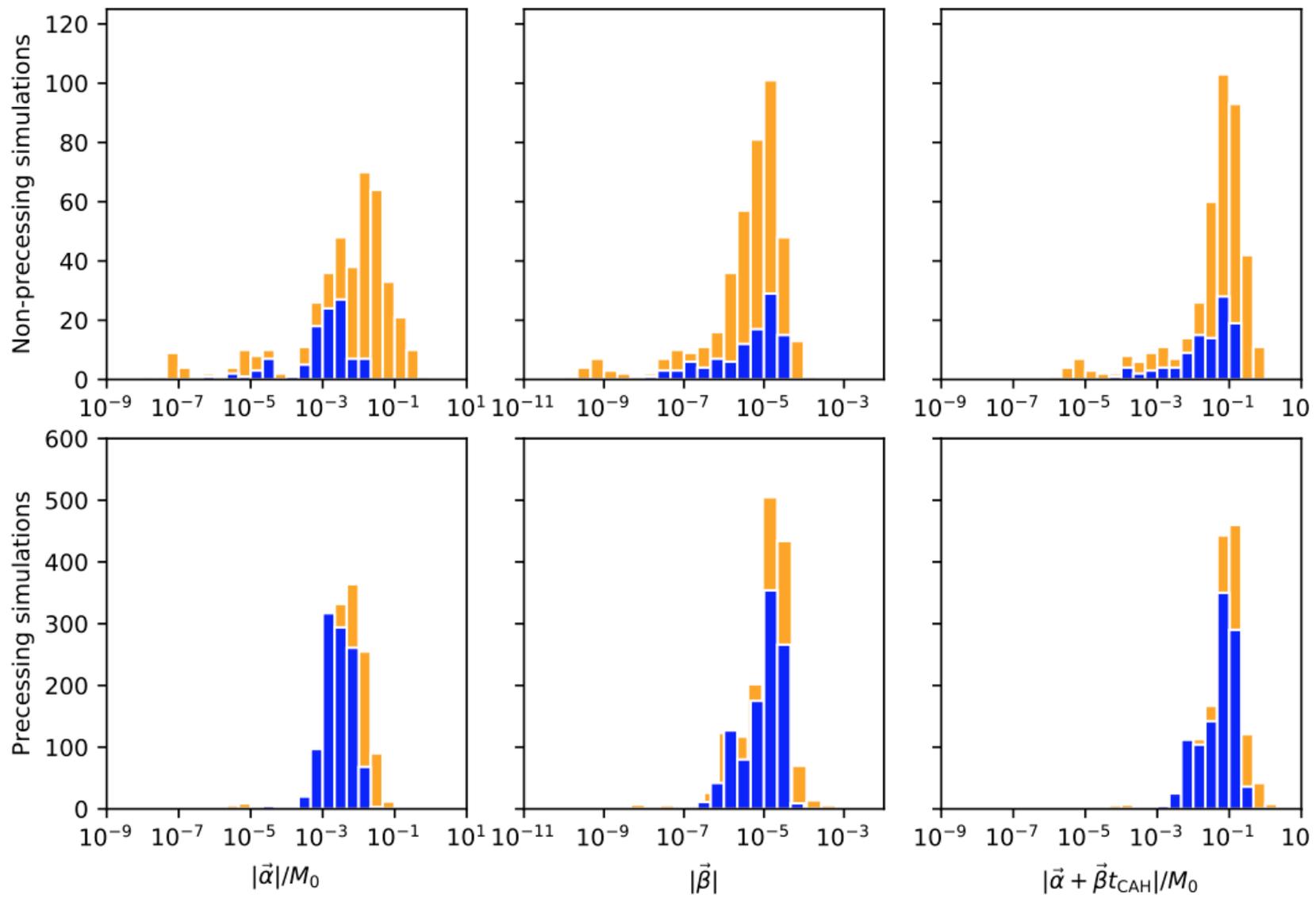
- Extrapolated waveform is  $f_{(0)}(u)$

# CoM correction

- "Gauge-invariant waveforms" 😬
- Displacements around 0.1M show up as big wiggles
- General BMS transformations allowed
- We just use boost and translation
- Even worse, we just use the horizon trajectories

<https://arxiv.org/abs/1509.00862> (<https://arxiv.org/abs/1509.00862>)

<https://arxiv.org/abs/1904.04842> (<https://arxiv.org/abs/1904.04842>)



# The problem

- Total catalog is ~16 TB
- Extrapolated waveform files currently average 95 MB each
- Users who want one waveform from each simulation need ~200 GB
- This is accelerating

# The solutions

- Eliminate redundant time data – factor of ~1.45 reduction
- Reduce the number of files – factor of ~8 reduction
  - Eliminate finite-radius files
  - Eliminate non-CoM files
- Store each file more efficiently – factor of ~4 reduction
  
- Total catalog size falls to ~500 GB
- Typical user needs ~10 GB

# Proposed spec

## Requirements

- H5 file
  - All datasets stored as XOR-ed UINT64s
  - Attribute format
  - Dataset time
    - Shape (n\_times, )
    - Must be strictly monotonic
  - Dataset modes
    - Shape (n\_times, 2\*n\_modes)
    - Conjugate-pair representation
    - Attribute e11\_min
    - Attribute e11\_max
  - Dataset log\_frame (optional)
    - Shape (0, 3), (1, 3), or (n\_times, 3)
- JSON file
  - same file name with .h5 → .json
  - more details on later slides

## Recommendations

- Corotating frame
- Truncation at  $10^{-10}$  of norm
- Add 0.0
- Compression options
- Choices of which files to distribute
- Directory structure and file names

# Corotating frame

- Defined as rotating frame that minimizes time dependence of modes
- Compute angular velocity of this frame by solving

$$\langle h | L L \cdot \vec{\omega} | h \rangle = -\langle h | L \partial_t | h \rangle$$

- The integrate the angular velocity to find a frame
- Constant of integration  $\rightarrow$  align  $z$  with dominant eigenvector of  $\langle h | L L | h \rangle$

<https://arxiv.org/abs/1302.2919> (<https://arxiv.org/abs/1302.2919>)

# Conjugate-pair mode representation

- $h_{\ell,m}$  and  $h_{\ell,-m}$  are at least somewhat redundant in corotating frame
- Define

$$s_{\ell,m} = \frac{h_{\ell,m} + \bar{h}_{\ell,-m}}{\sqrt{2}}$$

$$d_{\ell,m} = \frac{h_{\ell,m} - \bar{h}_{\ell,-m}}{\sqrt{2}}$$

- One of these will be zero (non-precessing systems) or small (otherwise), and thus compress well
- Store

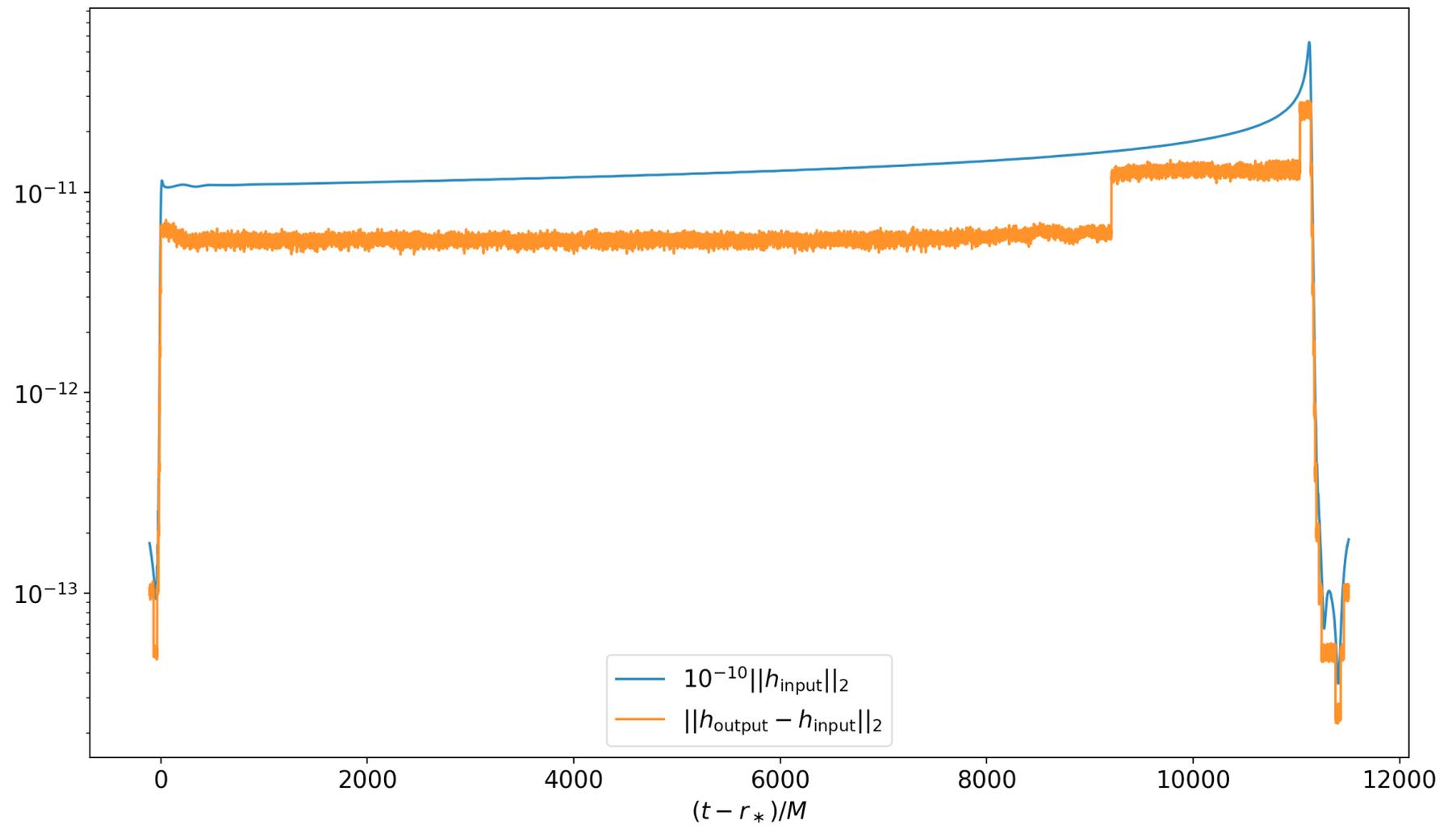
[ $d_{2,2}, d_{2,1}, h_{2,0}, s_{2,1}, s_{2,2}, d_{3,3}, d_{3,2}, d_{3,1}, h_{3,0}, s_{3,1}, s_{3,2}, s_{3,3}, \dots$ ]

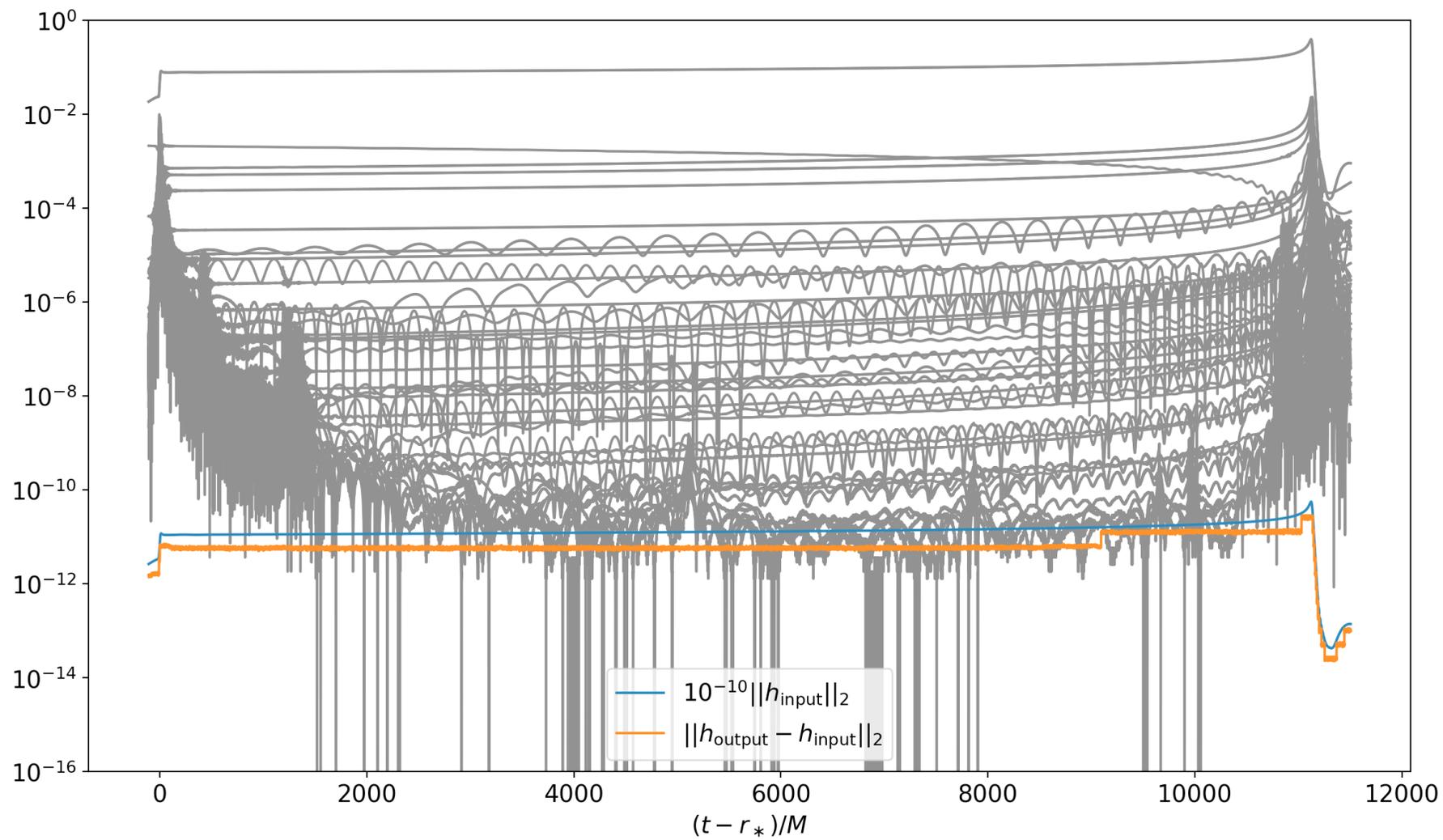
# Truncation

Zero out bits that affect norm at level below some chosen tolerance

- Only non-exact step
- Choose the *relative* tolerance `tol`
- Defines the *absolute* tolerance `tol * norm(data)` as function of time
- Drop bits at lower significance from conjugate-pair modes

```
def truncate(data, n_modes, tol=1e-10):  
    tol /= sqrt(n_modes) # Distribute ~evenly among all modes  
    abs_tol = tol * norm(data, axis=1)  
    power_of_2 = 2 ** (floor(-log2(abs_tol))) # 1 / (largest power of 2 <= abs_tol)  
    return round(data * power_of_2) / power_of_2
```





# Add 0.0

- Before truncation, many modes were small and noisy
- Signs were basically random
- After truncation, they are now  $\pm 0.0$
- The sign doesn't matter, but flips randomly, thus not compressing well
- Get rid of it by adding 0.0

# XOR with previous time step

- Bitwise difference between data at successive times
- For ~smooth data, we are left with few bits changing
- This compresses *very* well
- It's trivial to do, and to undo:

# Shuffle

- Bytes of similar significance more correlated than consecutive bytes
- Store bytes of given significance consecutively

```
In [5]: for n in np.random.rand(8): print(f"{n.view(np.uint64):064b}\n")
```

```
00111111110101110111101011110000011001010010000100010000111110010  
0011111111011000011100100100011101111111110100111010110100000100  
0011111111010000001001011010100110110110010011010100010100011000  
001111110101000001101100001100010100110111001011101111000000000  
0011111111010110101000011110111001110101100110100001110001011000  
001111111110110011010111110000111111001110111011000111111111100  
0011111111100100111001111100110110110101000110011111111010001011  
0011111111100000111001001101010101010010110101000000001011100101
```

# Format

- Corotating frame
- Paired modes
- Truncate
- XOR
- Multishuffle
- Bzip2

Are you scared yet?

# Changing to a new format

- Pros:
  - Significant reduction in size
  - ~No cost to add insignificant modes
  - Easier for maintainers
  - Easier for users
- Cons:
  - Different is bad

# Solve the problem with code

- `GW Frames` – old and ugly
- `scri` - powerful but complicated
- `sxs` - new and beautiful